Lecture No. 5



Phase Space Representation. Ensemble of Particles, Emittance.

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From Single Particle to a Beam

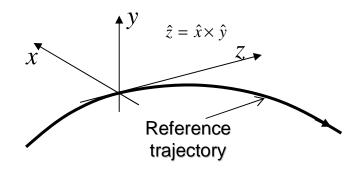


- The number of particles per bunch in most accelerators is typically included between 10⁵ to 10¹³.
- Integrating the particle motion for such a large number of particles along accelerators with length ranging from few meters up to tens of kilometers can prove to be a tough task.
- Fortunately, statistical mechanics gives us very developed tools for representing and dealing with sets of large number of particles.
- Quite often, the statistical approach can give us elegant and powerful insights on properties that could be hard to extract by approaching the set using single particle techniques.

A Convenient Reference Frame



- In accelerators we are interested in studying particles along their trajectory. A natural choice is to refer all the particles relatively to a reference trajectory.
- Such a trajectory is assumed to be the solution of the Lorentz equation for the particle with the nominal parameters (reference particle).
- In each point of this trajectory we can define (for example) a Cartesian frame moving with the reference particle.
- In this frame the reference particle is always at the origin and its momentum is always parallel to the direction of the z axis.



- The coordinates $\{x, y, z\}$ for an arbitrary particle represent its displacement relatively to the reference particle along the three directions.
- In the lab frame the particle moves on the curvilinear coordinate *s* with speed *dsldt*.

Phase Space Representation



In relativistic classical mechanics, the motion of a single particle is totally defined when, at a given instant t, the position r and the momentum p of the particle are given together with the forces acting on that position.

$$\overline{r}_{i} = x_{i} \, \hat{x} + y_{i} \, \hat{y} + z_{i} \, \hat{z}$$

$$\overline{F} = F_{x} \, \hat{x} + F_{y} \, \hat{y} + F_{z} \, \hat{z}$$

$$\overline{F} = F_{x} \, \hat{x} + F_{y} \, \hat{y} + F_{z} \, \hat{z}$$

It is quite convenient to use the so-called *phase space* representation, a 6-D space where the ith particle assumes the coordinates:

$$P_i \equiv \left\{ x_i, p_{xi}, y_i, p_{yi}, z_i, p_{zi} \right\}$$

In most accelerator physics calculations, the three planes can be considered with very good approximation as decoupled. In this situation, it is possible and convenient to study the particle evolution independently in each of the planes:

$$\left\{x_{i},p_{xi}\right\}$$

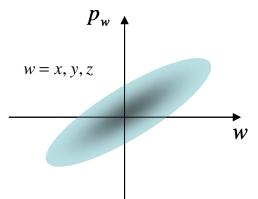
$$\left\{ y_{i},p_{yi}\right\}$$

$$\{z_i, p_{zi}\}$$

Particles Systems & Ensembles



The phase space can now be used for representing particles:



The set of possible states for a system of N particles is referred as *ensemble* in statistical mechanics.

In the statistical approach, the particles lose their individuality. The properties of the whole system as individual entity are now studied.

The system is fully represented by the density of particles f_{6D} and f_{2D} :

$$f_{6D}(x, p_x, y, p_y, z, p_z) dx dp_x dy dp_y dz dp_z$$
 $f_{2D}(w, p_w) dw dp_w$ $w = x, y, z$

The above expressions indicate the number of particles contained in the elementary volume of phase space for the 6D and 2D cases respectively.

$$\int f_{6D} dx dp_x dy dp_y dz dp_z = N \qquad \int f_{2D} dw dp_w = N \qquad w = x, y, z$$

Important properties of the density functions can now be derived.

Under particular circumstances, such properties allow to calculate the time evolution of the particle system without going through the integration of the motion for each single particle.

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Hamiltonian Systems



A system of variables q (generalized coordinates) and p (generalized momenta) is Hamiltonian when exists a function H(q, p, t) that allows to describe the evolution of the system by:

$$\frac{dq_i}{dt} = \frac{\partial H}{\partial p_i}$$

$$\frac{dp_i}{dt} = -\frac{\partial H}{\partial q_i}$$

$$\frac{dq_i}{dt} = \frac{\partial H}{\partial p_i}$$

$$\frac{dp_i}{dt} = -\frac{\partial H}{\partial q_i}$$

$$\overline{q} \equiv \{q_1, q_2, \dots, q_N,\}$$

$$\overline{p} \equiv \{p_1, p_2, \dots, p_N,\}$$

The function H is called Hamiltonian and q and p are referred as canonical conjugate variables.

In the particular case that q are the usual spatial coordinates $\{x, y, z\}$ and p their conjugate momenta $\{p_x, p_y, p_z\}$, H coincides with the total energy of the system:

$$H = U + T$$
 = Potential Energy + Kinetic Energy

Non-Hamiltonian Forces:

 Stochastic processes (collisions, quantum emission, diffusion, ...) •Inelastic processes (ionization, fusion, fission, annihilation, ...) Dissipative forces (viscosity, friction, ...) 6

The Continuity Equation



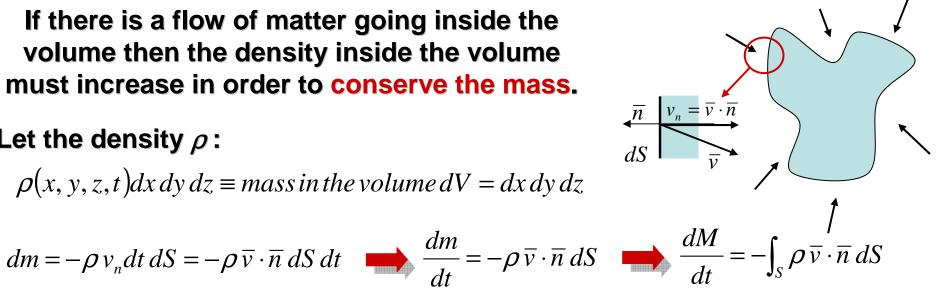
If there is a flow of matter going inside the volume then the density inside the volume must increase in order to conserve the mass.

Let the density ρ :

 $\rho(x, y, z, t)dx dy dz \equiv mass in the volume dV = dx dy dz$

$$dm = -\rho v_n dt dS = -\rho \overline{v} \cdot \overline{n} dS dt$$

$$\frac{dm}{dt} = -\rho \, \overline{v} \cdot \overline{n} \, dS$$



But it is also true that:

$$M = \int_{V} \rho \, dV \qquad \longrightarrow \qquad \frac{d}{dt} \int_{V} \rho \, dV = -\int_{S} \rho \, \overline{v} \cdot \overline{n} \, dS \qquad \longrightarrow \qquad \int_{S} \overline{F} \cdot \overline{n} \, dS = \int_{V} \nabla \cdot \overline{F} \, dV$$

$$\int_{S} \overline{F} \cdot \overline{n} \, dS = \int_{V} \nabla \cdot \overline{F} \, dV$$

$$\int_{S} \rho \, \overline{v} \cdot \overline{n} \, dS =$$

$$\int_{S} \rho \, \overline{v} \cdot \overline{n} \, dS = \int_{V} \nabla \cdot \rho \, \overline{v} \, dV \qquad \qquad \frac{d}{dt} \int_{V} \rho \, dV = -\int_{V} \nabla \cdot \rho \, \overline{v} \, dV$$

$$\frac{\partial f}{\partial x}$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \rho \, \overline{v} = 0$$

The continuity equation is a consequence of the consequence of the conservation law

The Liouville Theorem



Let us use the continuity equation with our phase space distributions. For simplicity we will use the 2D distribution, but the same exact results apply to the more general 6D case. $\frac{\partial \rho}{\partial t} + \nabla \cdot \rho \, \overline{v} = 0 \qquad \text{Let} \quad \rho \equiv f_{2D}(x, p_x) \text{ and } \overline{v} \equiv \{\dot{x}, \dot{p}_x\}$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \rho \, \overline{v} = 0$$

Let
$$\rho \equiv f_{2D}(x, p_x)$$
 and $\overline{v} \equiv \{\dot{x}, \dot{p}_x\}$



$$\frac{\partial f_{2D}}{\partial t} + \nabla \cdot f_{2D} \overline{v} = \frac{\partial f_{2D}}{\partial t} + \frac{\partial (\dot{x} f_{2D})}{\partial x} + \frac{\partial (\dot{p}_x f_{2D})}{\partial x} = \frac{\partial f_{2D}}{\partial t} + \frac{\partial f_{2D}}{\partial x} \dot{x} + \frac{\partial f_{2D}}{\partial p_x} \dot{p}_x + f_{2D} \frac{\partial \dot{p}_x}{\partial p_x} + f_{2D} \frac{\partial \dot{x}}{\partial x}$$

But our system is Hamiltonian $f_{2D} \frac{\partial \dot{x}}{\partial x} + f_{2D} \frac{\partial \dot{p}_x}{\partial p_x} = f_{2D} \frac{\partial^2 H}{\partial x \partial p_x} - f_{2D} \frac{\partial^2 H}{\partial x \partial p_x} = 0$



$$\frac{\partial f_{2D}}{\partial t} + \nabla \cdot f_{2D} \overline{v} = \frac{\partial f_{2D}}{\partial t} + \frac{\partial f_{2D}}{\partial x} \dot{x} + \frac{\partial f_{2D}}{\partial p_x} \dot{p}_x = \frac{df_{2D}}{dt}$$

$$\frac{df_{2D}}{dt} = 0$$

Liouville Theorem: The phase space density for a Hamiltonian system is an invariant of the motion. Or equivalently, the phase space volume occupied by the system is conserved.



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Decoupling the Problem: the Longitudinal Phase Space



- In most of existing accelerators the phase space planes are weakly coupled. In particular, we can treat the longitudinal plane independently from the transverse one in the large majority of the cases.
 - The effects of the weak coupling can be then investigated as a perturbation of the uncoupled case.
- In the longitudinal plane we apply our electric fields for accelerating the particles and changing their energy.
 - It becomes natural to use energy as one of the longitudinal plane variable together with its canonical conjugate time.
 - In accelerator physics, the *relative energy variation* δ and the *relative time 'distance'* τ with respect to a reference particle are often used:

$$\left| \delta = \frac{E - E_0}{E_0} \right|$$

$$\tau = t - t_0$$

- According to Liouville, in the presence of Hamiltonian forces, the area occupied by the beam in the longitudinal phase space is conserved.
 - More in Lecture 8......

The Transverse Phase Space



For the transverse planes $\{x, p_x\}$ and $\{y, p_y\}$, it is usually used a slightly modified phase space where the momentum components are replaced by:

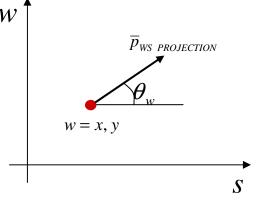
$$p_{xi} \to x' = \frac{dx}{ds}$$

$$p_{yi} \to y' = \frac{dy}{ds}$$

The physical meaning of the new variables:

$$x' = \frac{dx}{ds} = \tan \theta_x$$
 $y' = \frac{dy}{ds} = \tan \theta_y$

The relation between this new variables and the momentum (when $B_r = 0$) is:



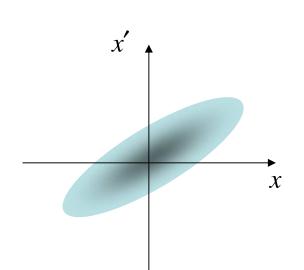
$$p_{x} = \gamma m_{0} \frac{dx}{dt} = \gamma m_{0} v_{s} \frac{dx}{ds} = \gamma \beta m_{0} c x'$$

$$p_{y} = \gamma \beta m_{0} c y'$$
where $\beta = \frac{v_{s}}{c}$ and $\gamma = (1 - \beta^{2})^{-1/2}$

Note that x and p_x are canonical conjugate variables while x and x' are not unless there is no acceleration (γ and β constant)

Definition of Emittance





We will consider the decoupled case and use the $\{w, w'\}$ plane where w can be either x or y.

We define as emittance the phase space area occupied by the system of particles, divided by π

$$\varepsilon_{w} = \frac{A_{ww'}}{\pi} \qquad w = x, y$$

As we previously shown, x and y are conjugate to x and y when $B_z = 0$ and in absence of acceleration. In this case, we can immediately apply the Liouville theorem and state that for such a system the emittance is an invariant of the motion.

This specific case is actually extremely important.

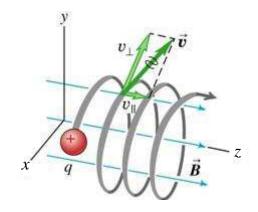
In fact, for most of the elements in a beam transferline, such as dipoles, quadrupoles, sextupoles, ..., the above conditions apply and the emittance is conserved.

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Emittance Conservation in The Presence of B_z



• When the B_z component of the magnetic field is present (solenoidal lenses for example), the transverse planes become coupled and the phase space area occupied by the system in each of the transverse planes is not conserved anymore.



- Anyway in this situation, the Liouville theorem still applies to the 4D transverse phase space where the ipervolume occupied by our system is still a motion invariant.
- Actually, if we rotate the spatial reference frame around the z axis by the Larmor frequency $\omega_L = qB_z/2\gamma m_0$, then the planes become decoupled and the phase space area in each of the planes is conserved again.

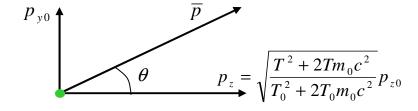
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Emittance Conservation in The Presence of Acceleration



When the particles in a beam undergo to acceleration, β and γ change and the variables x and x' are not canonical anymore. Liouville theorem does not apply and the emittance is not conserved.





$$y_0' = \tan \theta_0 = \frac{p_{y0}}{p_{z0}} = \frac{p_{y0}}{\beta_0 \gamma_0 m_0 c}$$

$$y' = \tan \theta = \frac{p_y}{p_z} = \frac{p_{y0}}{\beta \gamma m_0 c}$$

$$\frac{y'}{y_0'} = \frac{\beta_0 \, \gamma_0}{\beta \, \gamma}$$

 $y_0' = \tan \theta_0 = \frac{p_{y0}}{p_{z0}} = \frac{p_{y0}}{\beta_0 \gamma_0 m_0 c} \qquad y' = \tan \theta = \frac{p_y}{p_z} = \frac{p_{y0}}{\beta \gamma m_0 c} \qquad \frac{y'}{y_0'} = \frac{\beta_0 \gamma_0}{\beta \gamma}$ It can be shown that in this case $\frac{\mathcal{E}_y}{\mathcal{E}_{y0}} = \frac{y'}{y_0'} \qquad \boxed{\beta \gamma \mathcal{E}_y = \beta_0 \gamma_0 \mathcal{E}_{y0}}$

$$\beta \gamma \varepsilon_{y} = \beta_{0} \gamma_{0} \varepsilon_{y0}$$

The last expression tells us that the quantity $\beta \gamma \epsilon$ is a system invariant during acceleration. By defining the normalized emittance:

$$\varepsilon_{nw} = \beta \gamma \varepsilon_{w} \qquad w = x, y$$

We can say that the normalized emittance is conserved during acceleration. The acceleration couples the longitudinal plane with the transverse one: 13 the 6D emittance is still conserved but the transverse is not.

The Real Beam Case: r.m.s Emittance



For a real beam composed by N particles we can calculate the second order statistical moments of their phase space distribution:

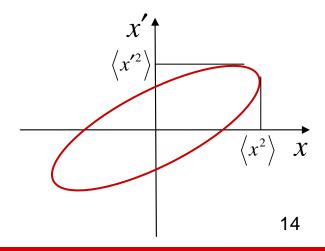
$$\langle x^{2} \rangle = \frac{\sum_{n=1}^{N} x_{n}^{2}}{N} \cong \frac{\int x^{2} f_{2D}(x, x') dx dx'}{\int f_{2D}(x, x') dx dx'} \qquad \langle x'^{2} \rangle = \frac{\sum_{n=1}^{N} x_{n}'^{2}}{N} \cong \frac{\int x'^{2} f_{2D}(x, x') dx dx'}{\int f_{2D}(x, x') dx dx'}$$
$$\langle x x' \rangle = \frac{\sum_{n=1}^{N} x_{n} x_{n}'}{N} \cong \frac{\int x x' f_{2D}(x, x') dx dx'}{\int f_{2D}(x, x') dx dx'}$$

And define the rms emittance as the quantity: $|\varepsilon_{rms} = \sqrt{\langle x^2 \rangle \langle x'^2 \rangle} - \langle x x' \rangle^2$

$$\varepsilon_{rms} = \sqrt{\langle x^2 \rangle \langle x'^2 \rangle - \langle x x' \rangle^2}$$

This is equivalent to associate to the real beam an equivalent or phase ellipse in the phase space with area $\pi \varepsilon_{rms}$ and equation:

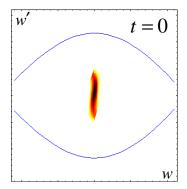
$$\frac{\left\langle x^{\prime 2} \right\rangle}{\varepsilon_{rms}} x^{2} + \frac{\left\langle x^{2} \right\rangle}{\varepsilon_{rms}} x^{\prime 2} - 2 \frac{\left\langle x \ x^{\prime} \right\rangle}{\varepsilon_{rms}} x \ x^{\prime} = \varepsilon_{rms}$$

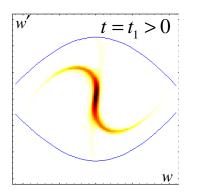


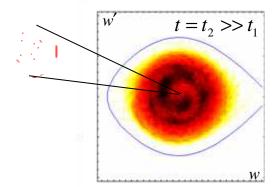
Nonlinear Forces and Filamentation



- In the case of a Hamiltonian system, as a consequence of the Liouville Theorem the emittance is conserved
 - This is true even when the forces acting are on the system are nonlinear (space charge, nonlinear magnetic and/or electric fileds, ...)
 - This is not true in the case of the rms emittance.
- •In the presence of nonlinear forces the rms emittance is not conserved
- Example: filamentation. Particles with different phase space coordinates, because of the nonlinear forces, move with different phase space velocity







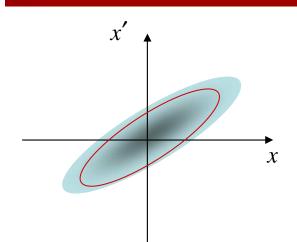
• The emittance according to Liouville is still conserved.

But the rms emittance calculated for increasing times increases.15

The Twiss Parameters



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We saw that a beam with arbitrary phase space distribution can be represented by an equivalent ellipse with area equal to the rms emittance divided by π and with equation:

$$\frac{\left\langle w'^{2}\right\rangle}{\varepsilon_{w \, rms}} w^{2} + \frac{\left\langle w^{2}\right\rangle}{\varepsilon_{w}} w'^{2} - 2 \frac{\left\langle w \, w'\right\rangle}{\varepsilon_{w}} w \, w' = \varepsilon_{w} \qquad w = x, y$$

A convenient representation for this ellipse, often used in accelerator physics, is the one by the so-called *Twiss Parameters* β_T , γ_T and α_T :

$$\beta_{Tw}w'^2 + \gamma_{Tw}w^2 + 2\alpha_{Tw}ww' = \varepsilon_w$$
 $w = x, y$ with $\beta_{Tw}\gamma_{Tw} - \alpha_{Tw}^2 = 1$

with
$$\beta_{Tw} \gamma_{Tw} - \alpha_{Tw}^2 = 1$$

The status of the beam at a given moment is totally defined when the emittance and two of the Twiss parameters are known.

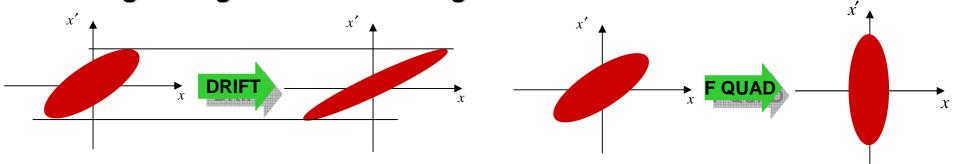
By comparing the two ellipse equations, we can derive:

$$\langle w^2 \rangle = \beta_{Tw} \varepsilon_w \qquad \langle w'^2 \rangle = \gamma_{Tw} \varepsilon_w \qquad \langle w w' \rangle = -\alpha_{Tw} \varepsilon_w \qquad w = x, y$$

& Emittance Propagating the Twiss Parameters



When the beam propagates along the beamline, the eccentricity and the orientation of the equivalent ellipse change while the area remains constant (Liouville theorem). In other words, the Twiss parameters change along the line according to the action of the line elements.



The single particle matrix formalism can now be extended to the Twiss parameters. For example for a drift of length L in the horizontal plane:

$$\begin{pmatrix} x \\ x' \end{pmatrix} = \begin{pmatrix} 1 & L \\ 0 & 1 \end{pmatrix} \begin{pmatrix} x_0 \\ x'_0 \end{pmatrix} \Rightarrow \begin{cases} x = x_0 + Lx'_0 \\ x' = x'_0 \end{cases}$$

$$\beta_{T} \mathscr{L} = \beta_{T0} \mathscr{L} + L^{2} \gamma_{T0} \mathscr{L} - 2L \alpha_{T0} \mathscr{L}$$

$$\Rightarrow \gamma_{T} \mathscr{L} = \gamma_{T0} \mathscr{L}$$

$$-\alpha_{T} \mathscr{L} = L \gamma_{T0} \mathscr{L} - \alpha_{T0} \mathscr{L}$$

$$\langle x^{2} \rangle = \langle (x_{0} + Lx'_{0})^{2} \rangle = \langle x_{0}^{2} \rangle + L^{2} \langle x'_{0}^{2} \rangle + 2L \langle x_{0}x'_{0} \rangle$$

$$\Rightarrow \langle x'^{2} \rangle = \langle x'_{0}^{2} \rangle$$

$$\langle xx' \rangle = \langle (x_{0} + Lx'_{0})x'_{0} \rangle = L \langle x'_{0}^{2} \rangle + \langle x_{0}x'_{0} \rangle$$

$$\Rightarrow \begin{pmatrix} \boldsymbol{\beta}_T \\ \boldsymbol{\gamma}_T \\ \boldsymbol{\alpha}_T \end{pmatrix} = \begin{pmatrix} 1 & L^2 & -2L \\ 0 & 1 & 0 \\ 0 & -L & 1 \end{pmatrix} \begin{pmatrix} \boldsymbol{\beta}_{T0} \\ \boldsymbol{\gamma}_{T0} \\ \boldsymbol{\alpha}_{T0} \end{pmatrix}$$

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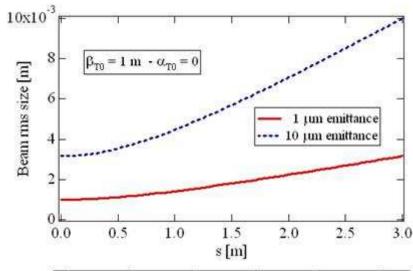
Getting Familiar with the Concept of Emittance



A couple of examples:

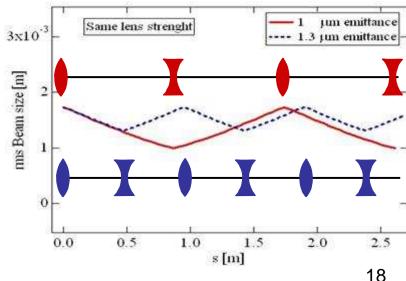
Propagation of beams with different emittance through a drift space





Propagation of beams with different emittance through a FODO lattice

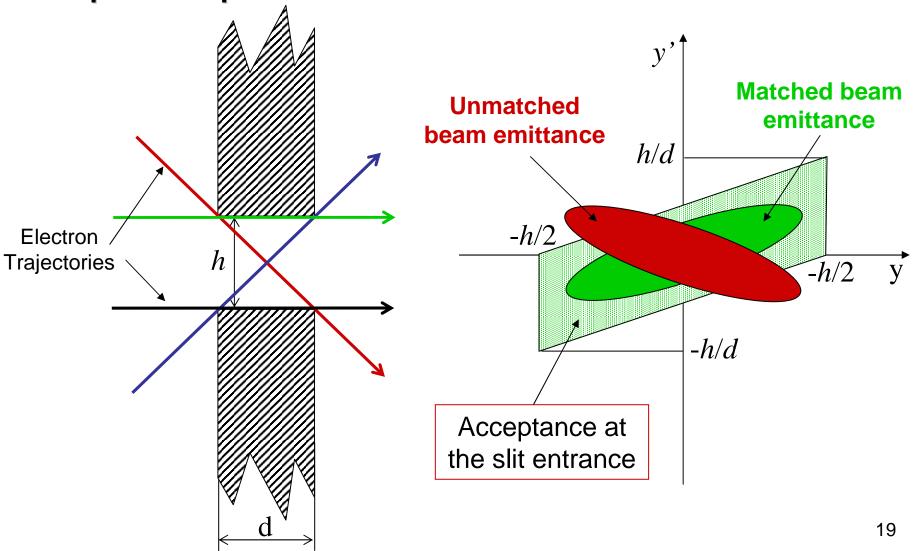




The Concept of Acceptance



Example: Acceptance of a slit



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The Emittance Role In Accelerator Applications



The emittance is a fundamental parameter in most accelerator applications.

- Electron microscopes: High resolution requires lower emittances
- Free electron lasers (FEL): Intensity of the radiation strongly depends on emittance. The smaller the better
- Synchrotron light sources: smaller emittances gives higher brightness
- Colliders: higher emittances give higher luminosity (in beam-beam limited regime)

• . . .

Possible Homework



• Prove the relation $d\varepsilon_0 = y'ly'_0$, where ε and ε_0 are the vertical emittance after and before acceleration by a field E_z , and y' and y'_0 are the divergences after and before acceleration. Tip: use the definition of rms emittance

• Calculate the Twiss parameter transport matrix for both planes of a focusing quadrupole in the thin lens approximation.